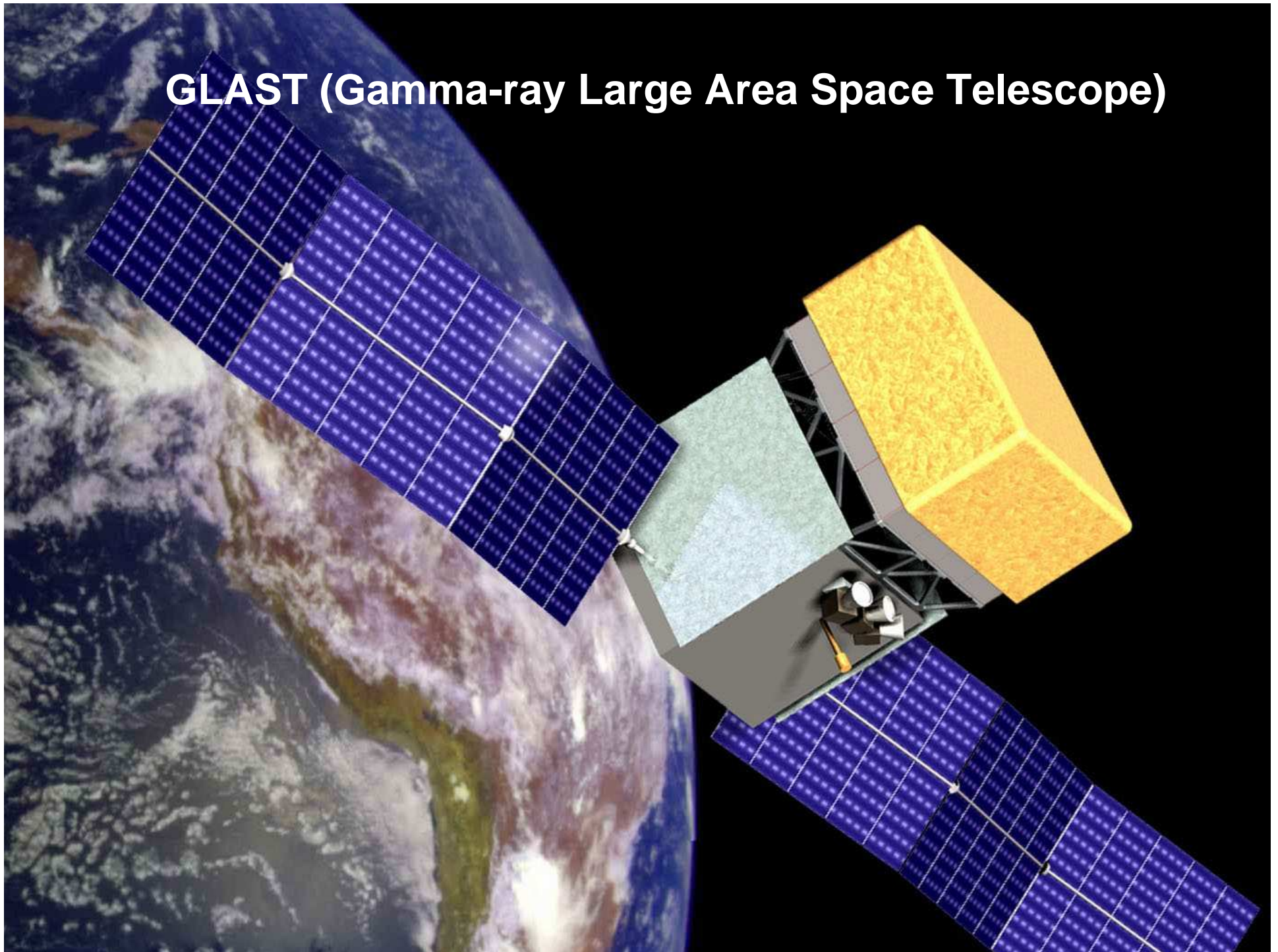
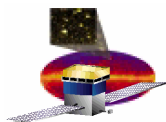


GLAST (Gamma-ray Large Area Space Telescope)





LAT (Large Area Telescope) 30 MeV-300 GeV

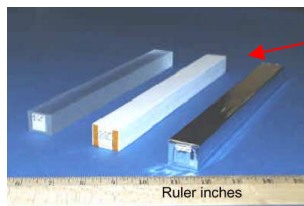
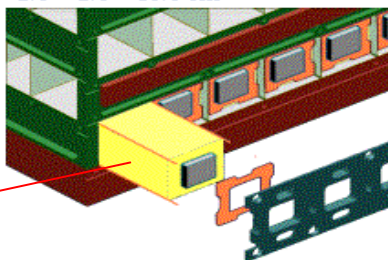
Si-W tracker

pitch = 201 μm
 $12 \times 2.5\% X_0$
 $+ 4 \times 25\% X_0$



CsI Calorimeter

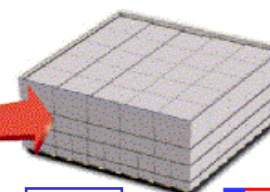
$8.6 X_0$ 8×12 bars
 $2.0 \times 2.8 \times 35.1$ cm



Pair conversion telescope

16 towers

- Veto
- Tracker
- Calorimeter



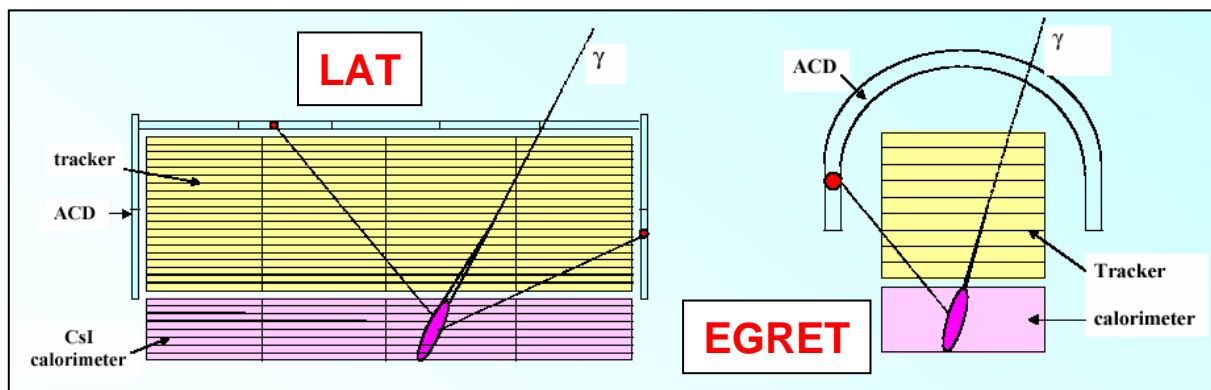
ACD

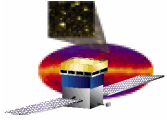
scintillator tiles
 0.9997 efficiency



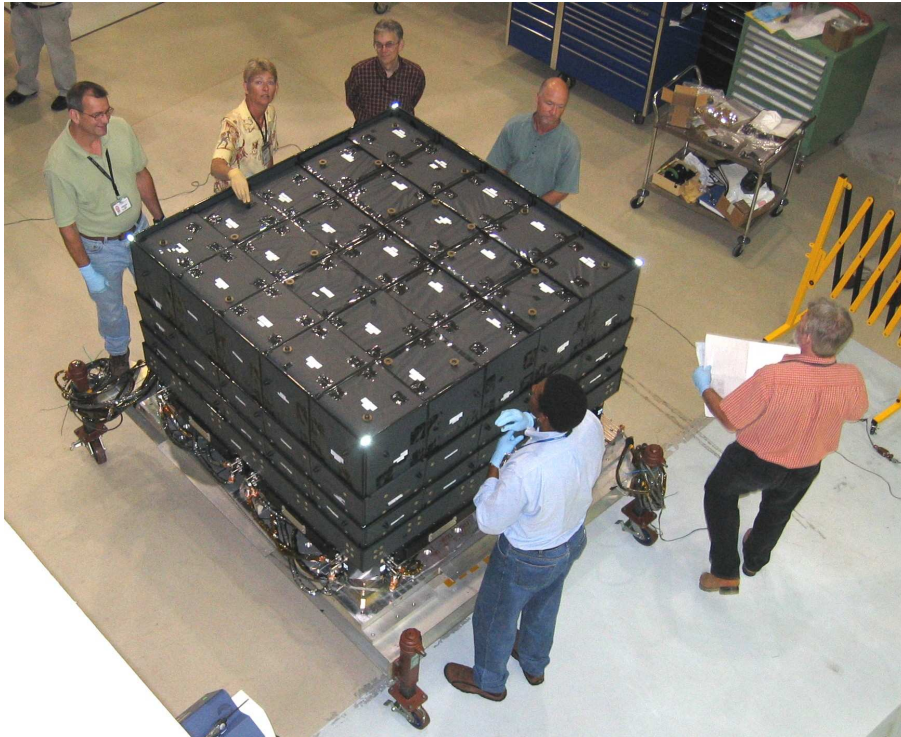
Science

Pulsars, SNRs, Blazars,
 GRBs, Dark Matter,
 Diffuse emissions





Anti-Coincidence Detector Complete

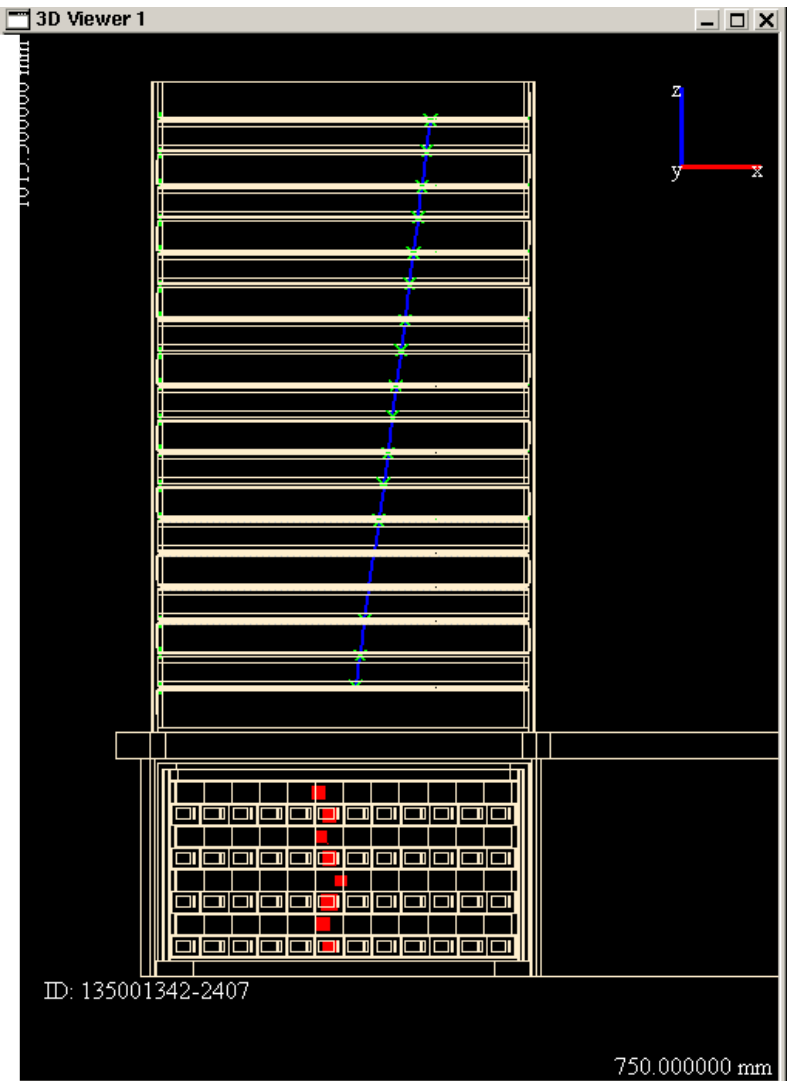


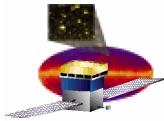
ACD before installation of
Micrometeoroid Shield



ACD with Micrometeoroid Shield
and Multi-Layer Insulation (but
without Germanium Kapton outer
layer)

David J. Thompson, GSFC
Thomas E. Johnson, GSFC





Program at SPS

Electrons used for testing the response to high-energy EM showers:

- energy reconstruction
- “backsplash” (firing of ACD tiles by backward–emitted electrons)
- behavior of electronics at high-rate....

GEANT4-based Monte-Carlo simulations will be tuned accordingly.

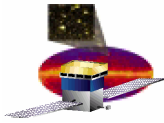
Highest energy, low background required

Protons needed for benchmarking the MC simulations developed for cosmic-ray background rejection.

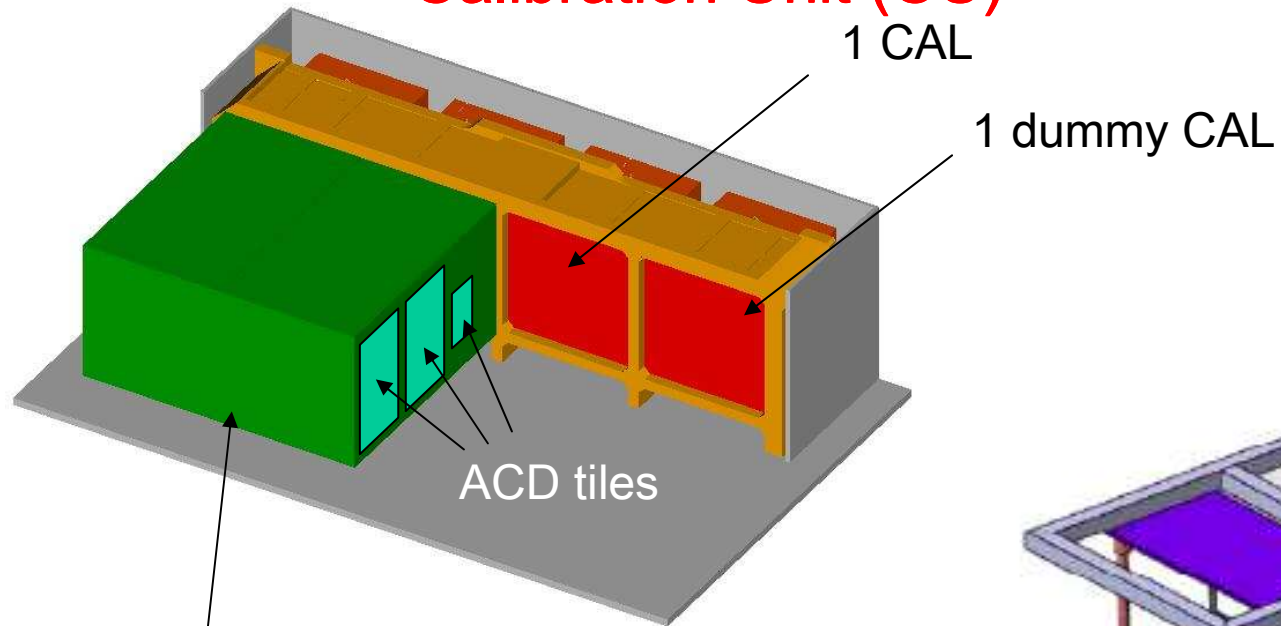
Low e background required

Particle	Energy (GeV)	Angles (deg.)	#Positions per angle	Estimated Statistics	Trigger	Configuration
Electrons	10, 20, 50, 100, 200, 300	0, 20, 40, 60	12	200k	Ext. *	Flight
	100	90	2	200k	Ext.	Flight
	10, 20, 50, 100, 200, 300	180	2	200k	Ext.	Flight
hadrons	10,20,50,100	0, 90	2	1 M	Ext.	Flight

* External trigger from plastic scintillator detectors in the beam line

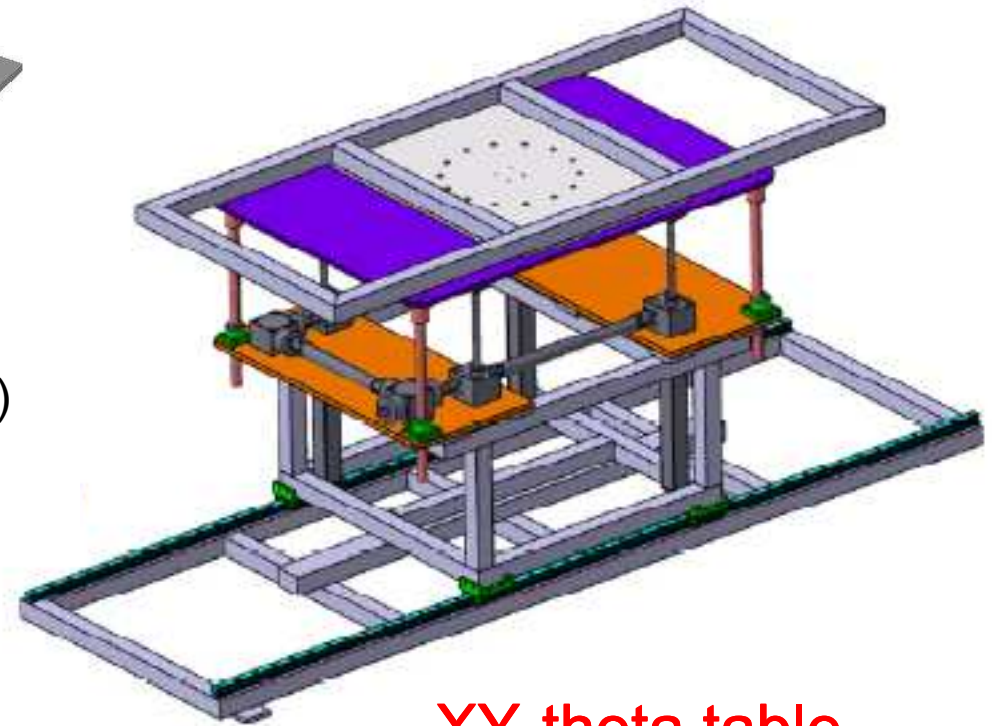


Calibration Unit (CU)

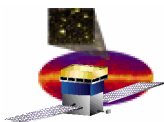


Two flight-spare "towers" (Tracker+CAL)
enclosed in a 2 mm-thick container

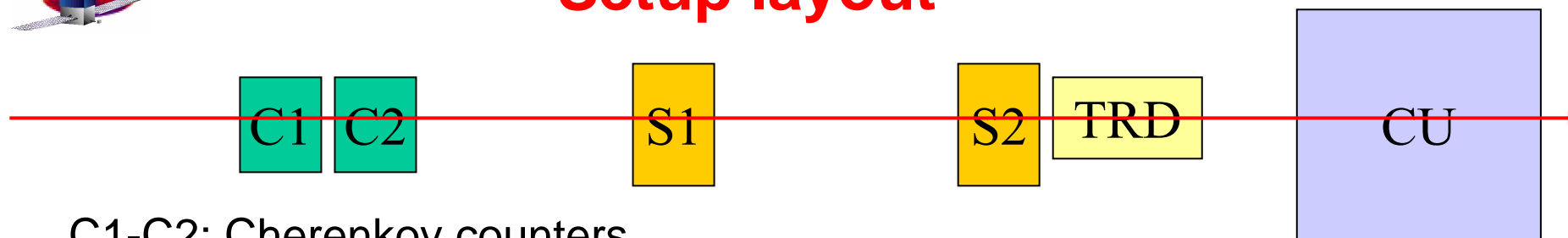
Detector at CERN



XY-theta table



Setup layout



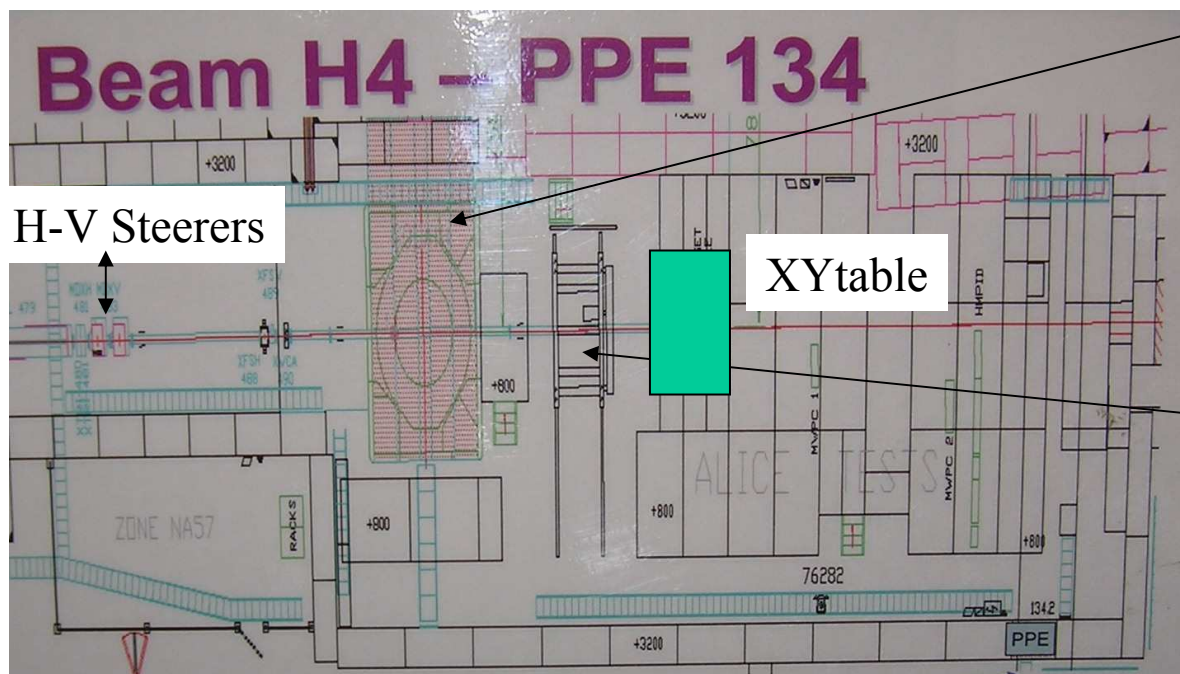
C1-C2: Cherenkov counters

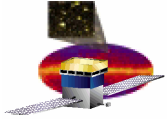
S1-S2: trigger scintillators

TRD

distance from end of vacuum beam pipe to table?

length of Cherenkov counters?





Beams

Rates: most cases $< 1\text{ kHz}$. A few runs with 10 kHz at the most favorable energy (100 GeV ?)

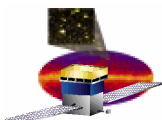
We strongly request the **0-mrad beam**:

- **perfect for electrons**: highest energy, low contamination, rate high enough (as found in GLAST run in 2002)
- possible use of **tertiary hadron beams** ($E < 100\text{ GeV}$) if rate is high enough.
Rates? Contamination?

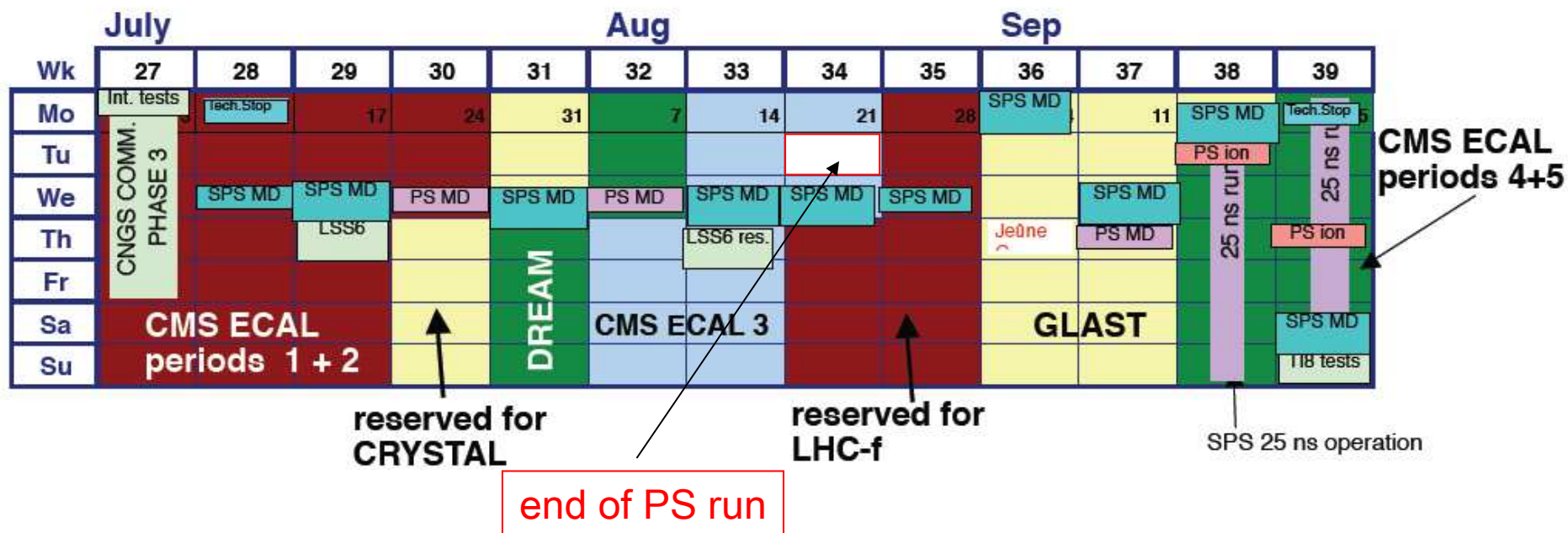
PID (in case the 0-mrad mode is not available)

For secondary hadron beams at low energy, we would use Cerenkov counters.
Use of N_2 or He_2 ? Can we switch from one to the other? Max.–Min. pressure?

Possible use of an ancillary calorimeter at high energy.



Planning

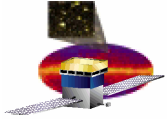


Beam: When do we know if there will be constraints on the beam due to H2 (CMS?)?

Installation: - access to the cave?

Barracks: - Will the two barracks be available?
- When can we move in? out?

Storage: clean room: availability, condition of use



Miscellaneous

installation:

X-Y table: position with respect to magnet
cables between cave and barracks: how many? how long?
ground in cave: via electric power (plugs)?
what signals to/from the control room?
use of He bag?

safety: course (when?, what?)
security visit: when?, who?, what?
GLIMOS
film badges

tunnel permit: still possible?

Name of responsible persons for: safety
crane+area
surveyor